



International
Centre for
Radio
Astronomy
Research

Identifying transient astronomical sources in MeerKAT light curves using Gaussian processes

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THE UNIVERSITY OF
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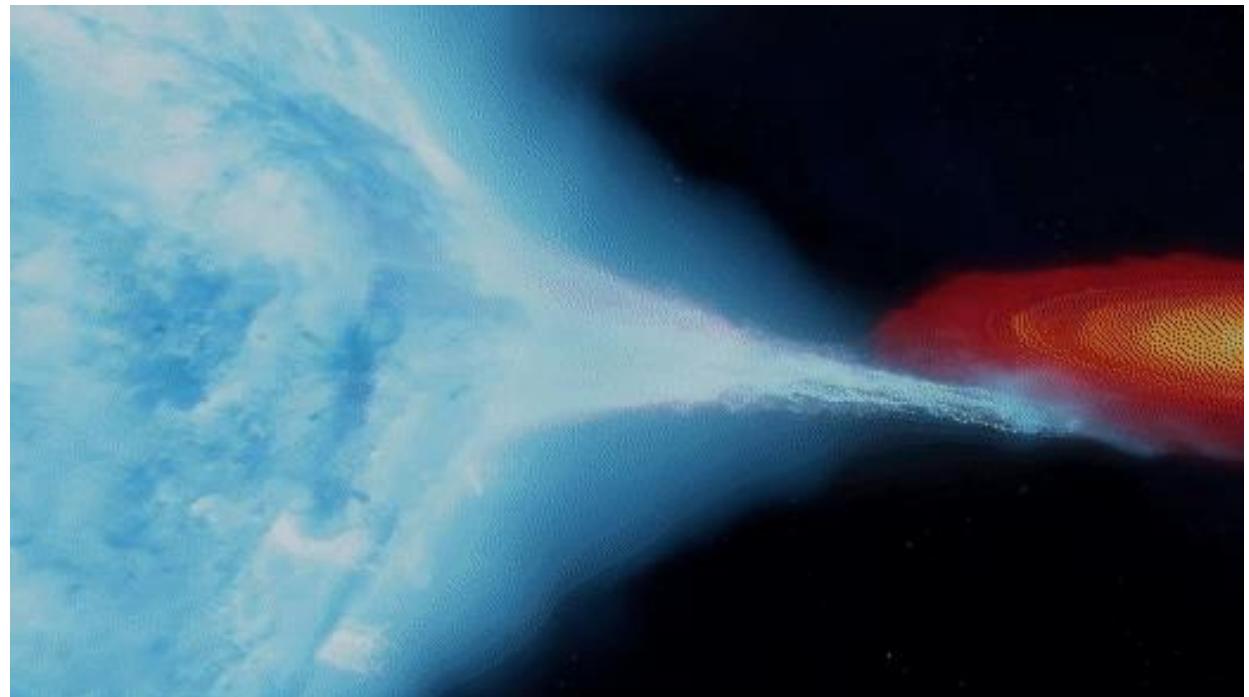


Government of Western Australia
Department of the Premier and Cabinet
Office of Science

Twinkle twinkle...

A *transient* is an astrophysical phenomenon whose brightness changes drastically over observable time.

- Supernovae
- Variable stars, e.g., pulsating,
• eclipsing binaries.
- Gamma-ray bursts (GRBs)
- Fast radio bursts (FRBs)
- Transiting planets
- Active galactic nuclei (AGN)
- Accreting blackholes
- and lots more...

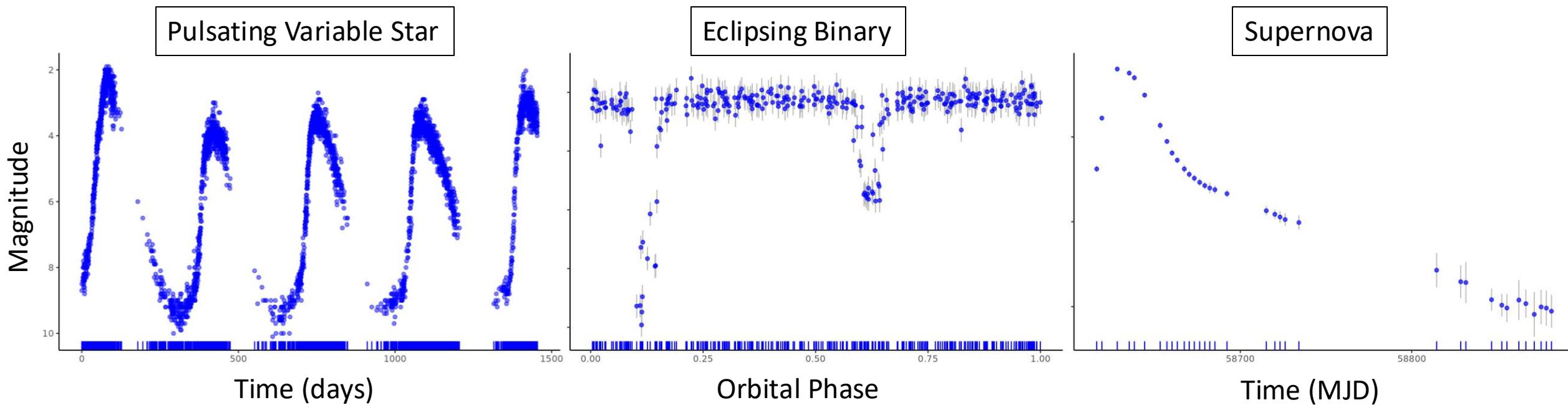


Artist's impression of the Cygnus X-1 system. Credit: ICRAR

Light Curves

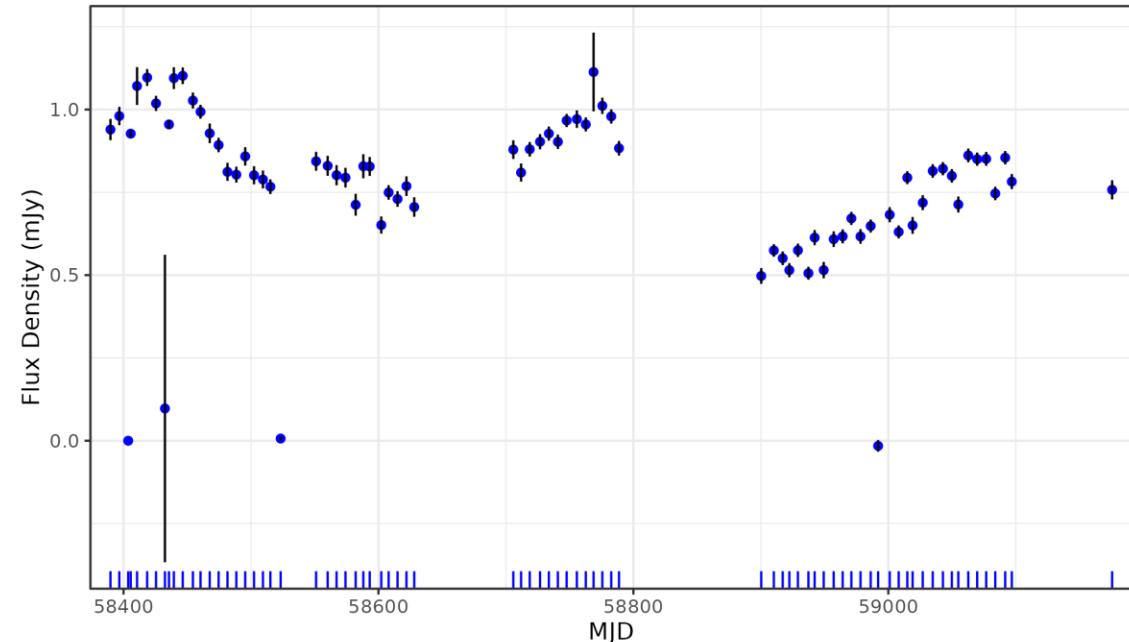
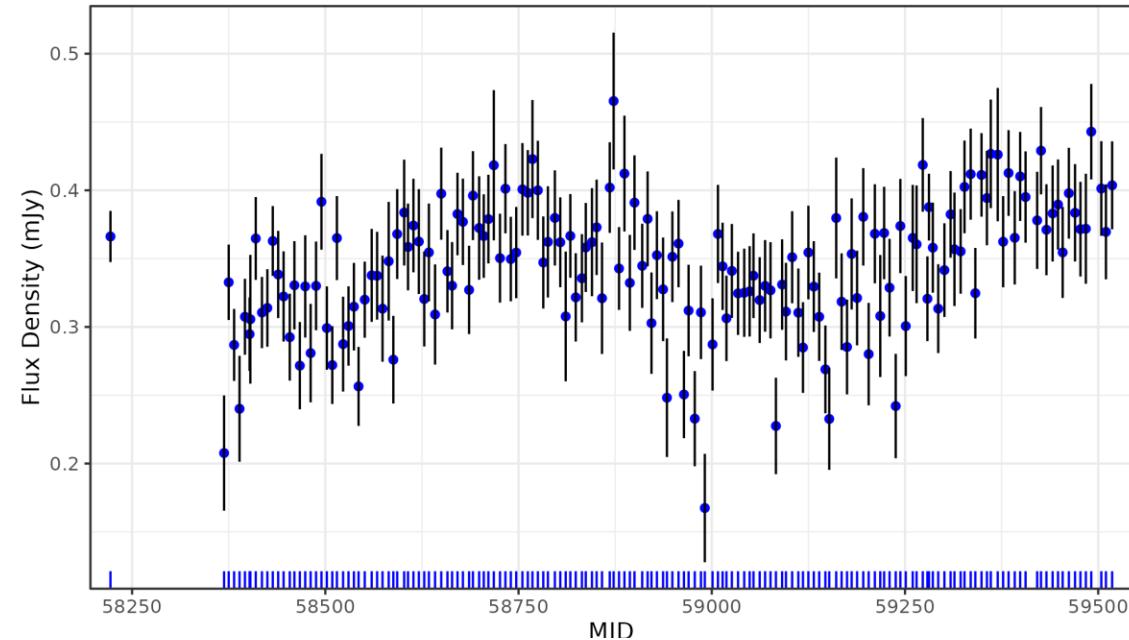
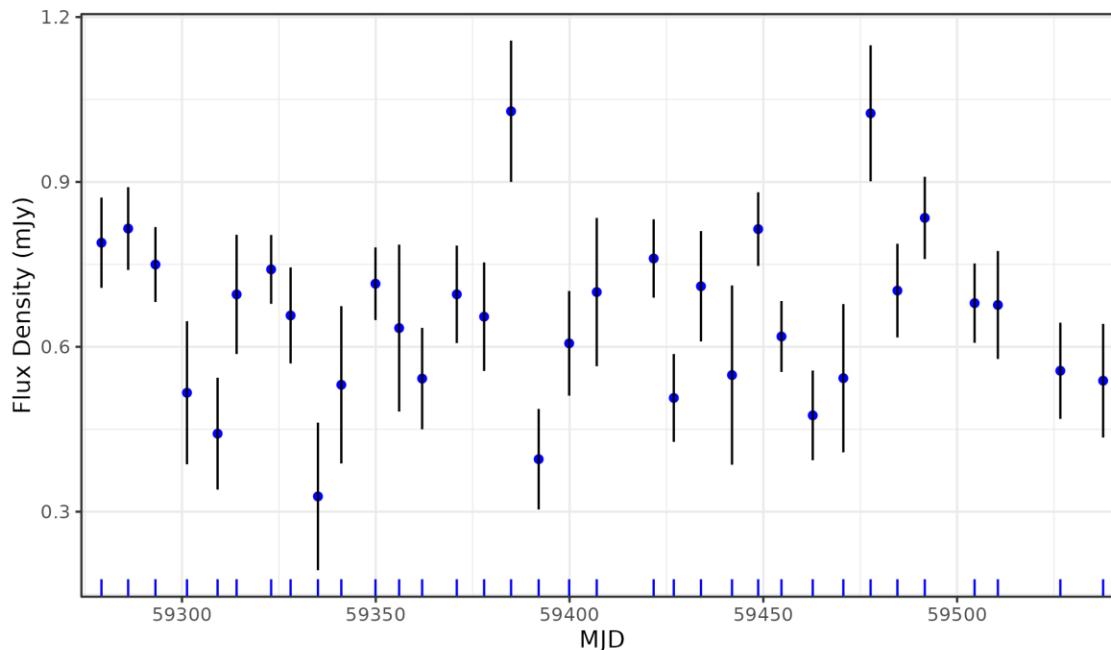
Light curves are time series describing the brightness of a source over time.

- The shape of a light curve can reveal the type of object or event.
- Variability in brightness can reveal information about the processes underlying the observed phenomenon.



Patchy Data

- Different cadences
- Sparse observations
- Irregular sampling
- Varying noise levels



Characterising Light Curves

Oversimplified

- Fewer parameters
- Scales easily
- High information loss

Overspecified

- Many parameters
- High discriminatory power
- Overfitting

Model light curves as a Gaussian Process (GP)



Gaussian Processes (GPs)

Extend multivariate Gaussian to ‘infinite’ dimensions.

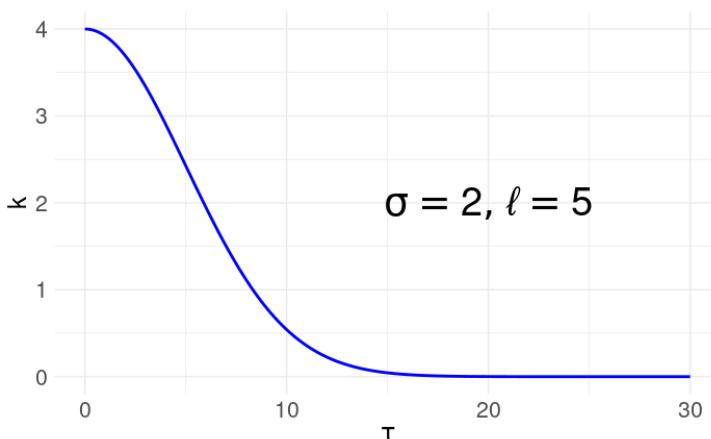
- Mean function, $\mu(t)$
- Covariance or **kernel function**, $\kappa(\mathbf{t}, \mathbf{t})$

$$\begin{bmatrix} Y_1 \\ Y_2 \\ \vdots \end{bmatrix} = \mathbf{Y} \sim GP(\mu(t), \Sigma)$$

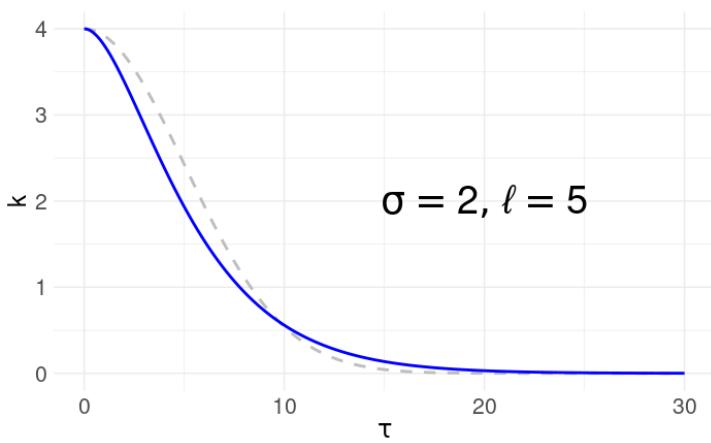
where $\mu = \mu(t_i)$ and $\Sigma_{ij} = \kappa(\mathbf{t}_i, \mathbf{t}_j)$, for $i, j = 1, 2, \dots$

Rather than specifying a fixed covariance matrix with fixed dimensions, compute covariances using the kernel function.

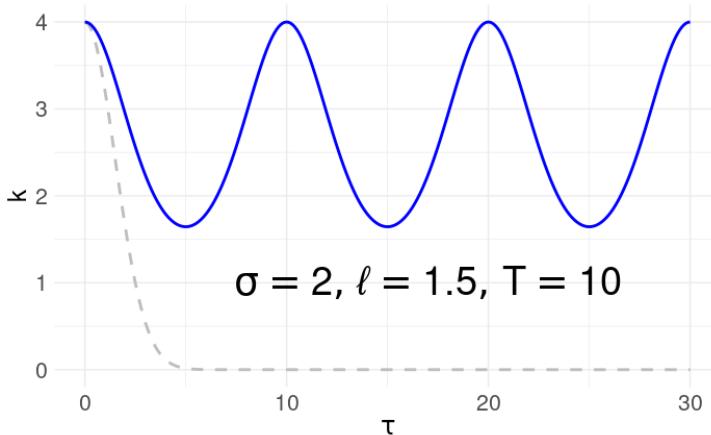
$$\tau = |t_r - t_c|; \sigma, \ell, T > 0$$



$$\kappa(\tau; \sigma, \ell) = \sigma^2 \exp\left\{-\frac{1}{2}\left(\frac{\tau}{\ell}\right)^2\right\} \quad \text{Squared Exponential}$$

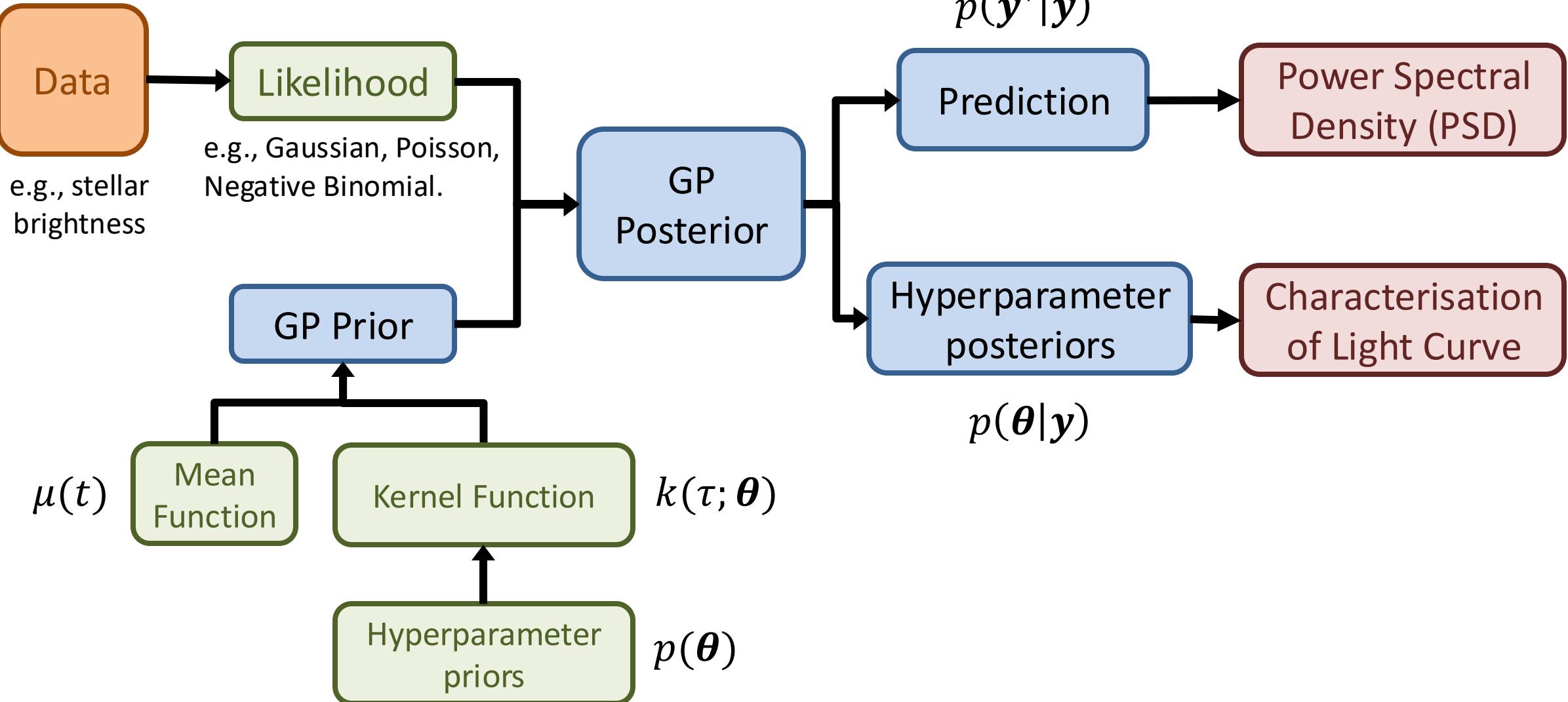


$$\kappa(\tau; \sigma, \ell) = \sigma^2 \left(1 + \sqrt{3}\frac{\tau}{\ell}\right) \exp\left\{-\sqrt{3}\frac{\tau}{\ell}\right\} \quad \text{Matern 3/2}$$



$$\kappa(\tau; \sigma, \ell, T) = \sigma^2 \exp\left\{-\frac{2}{\ell^2} \sin^2\left(\pi \frac{\tau}{T}\right)\right\} \quad \text{Periodic}$$

Modelling Workflow



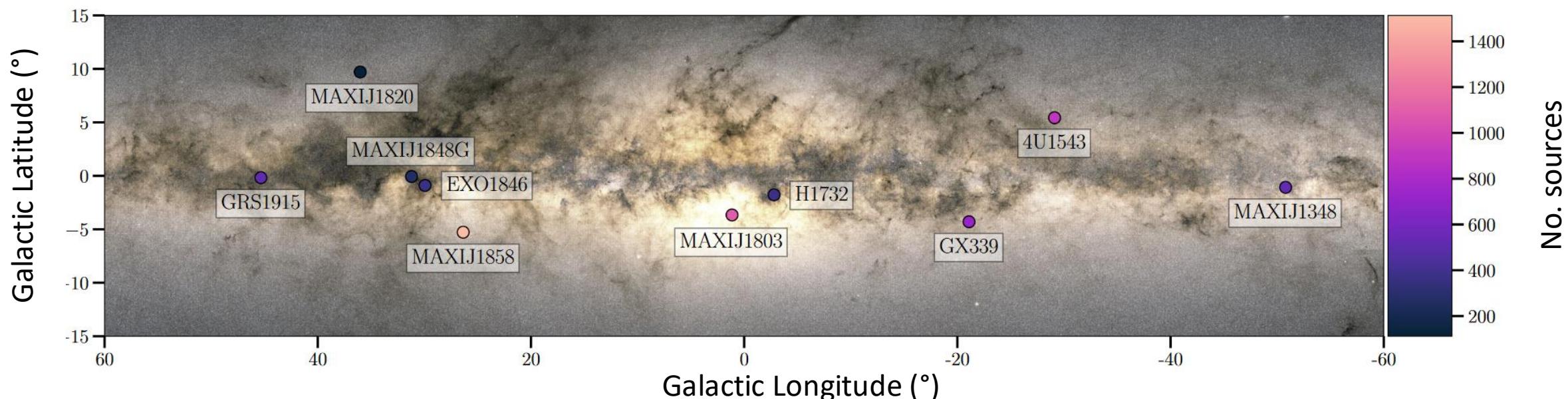


ThunderKAT Survey

- The HUNt for Dynamic and Explosive Radio transients with MeerKAT
- Field of view of ≈ 1 square degree
- 6,394 radio light curves over 10 fields
- Flux density measurements + standard errors



MeerKAT Radio Telescope (Credit: SARAO)



Gaussian Process Model

$$\mathbf{Y} \sim \mathcal{N}(f, \hat{\mathbf{e}}^2)$$

$$f \sim \text{GP}(\mathbf{0}, \mathbf{K}_{N \times N})$$

$$\mathbf{K}_{rc} = \kappa(t_r, t_c | \boldsymbol{\theta})$$

$$= \kappa_1(\tau; \sigma_{SE}, \ell_{SE})$$

Squared Exponential

$$+ \kappa_2(\tau; \sigma_{M32}, \ell_{M32})$$

Matern 3/2

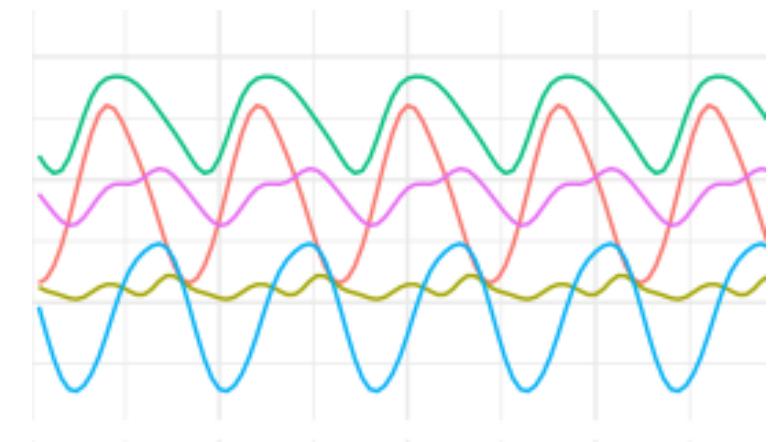
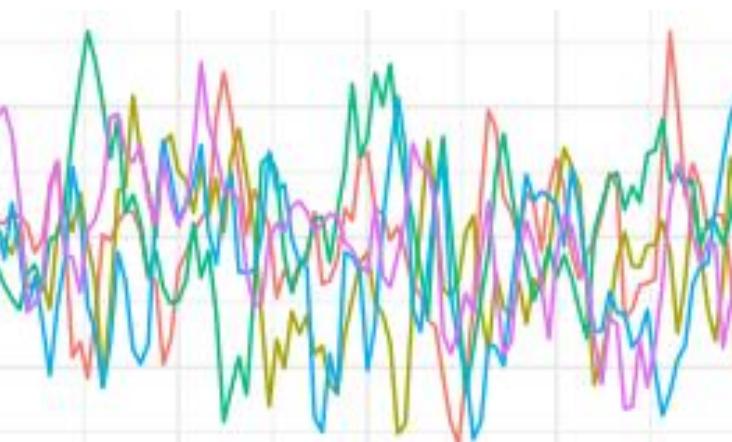
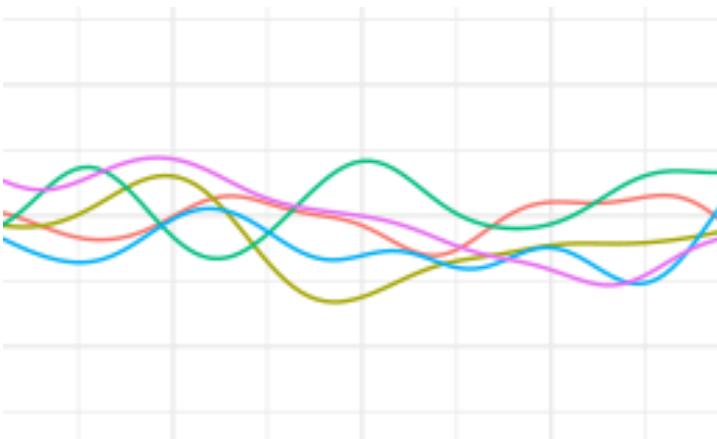
$$r, c = 1, \dots, N.$$

$$\boldsymbol{\theta} = (\sigma_{SE}, \ell_{SE}, \sigma_{M32}, \ell_{M32}, \sigma_P, \ell_P, T)$$

$$+ \kappa_3(\tau; \sigma_P, \ell_P, T)$$

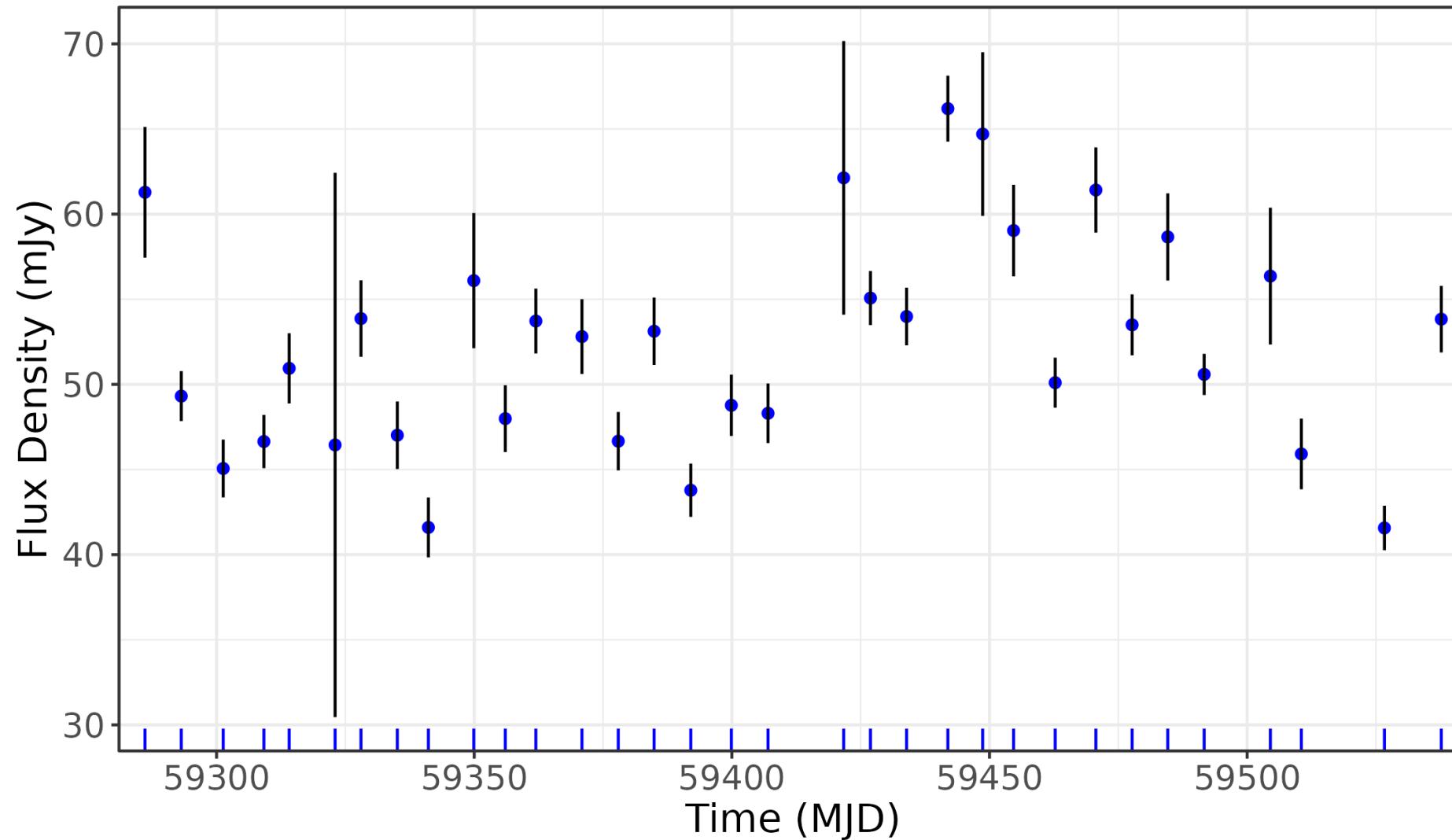
Periodic

Covariance
Kernel



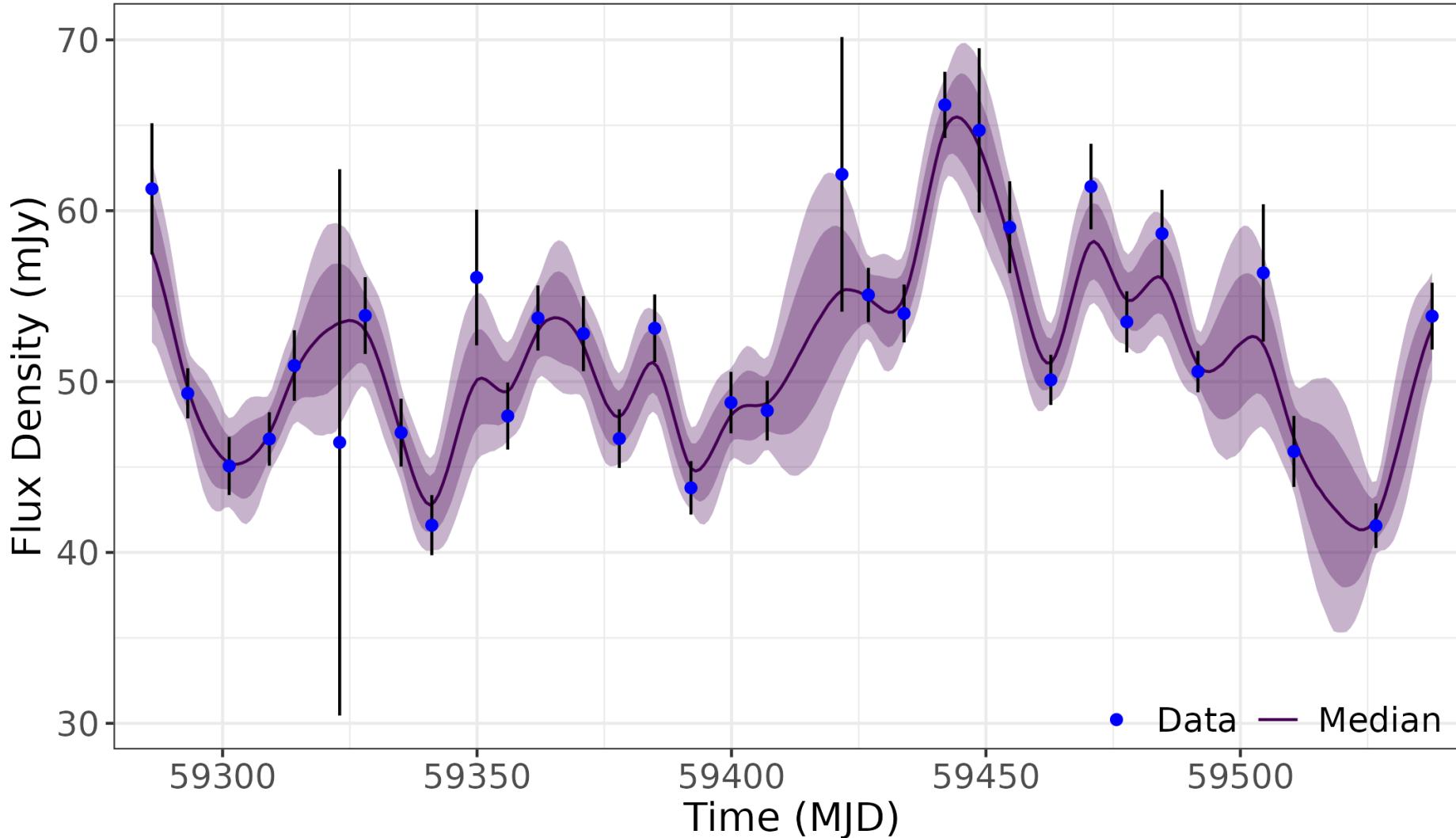
GP Fitting Example

N = 33, Duration = 215 days, Field = J1848G



Posterior Predictive Samples

N = 33, Duration = 215 days, Field = J1848G



$$\sigma_{\text{SE}} = 0.35$$

$$\ell_{\text{SE}} = 48.5$$

$$\sigma_{\text{M32}} = 1.20$$

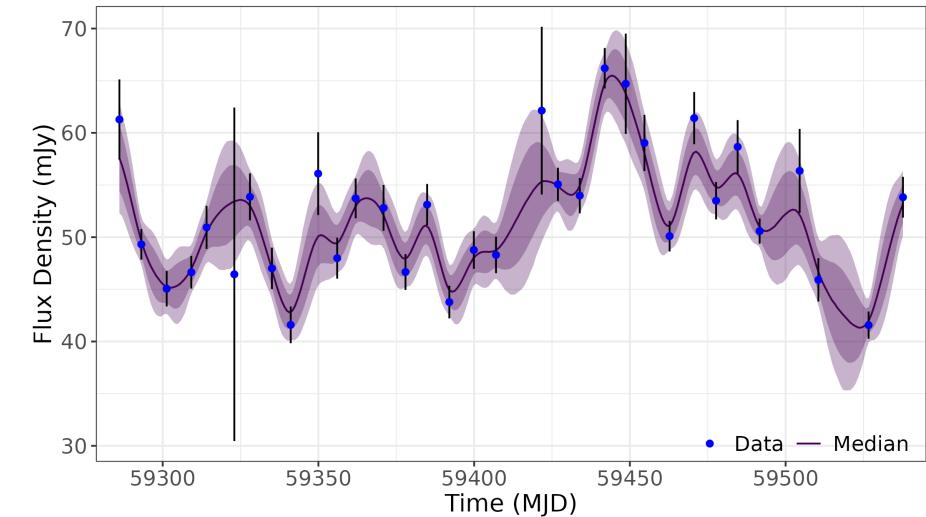
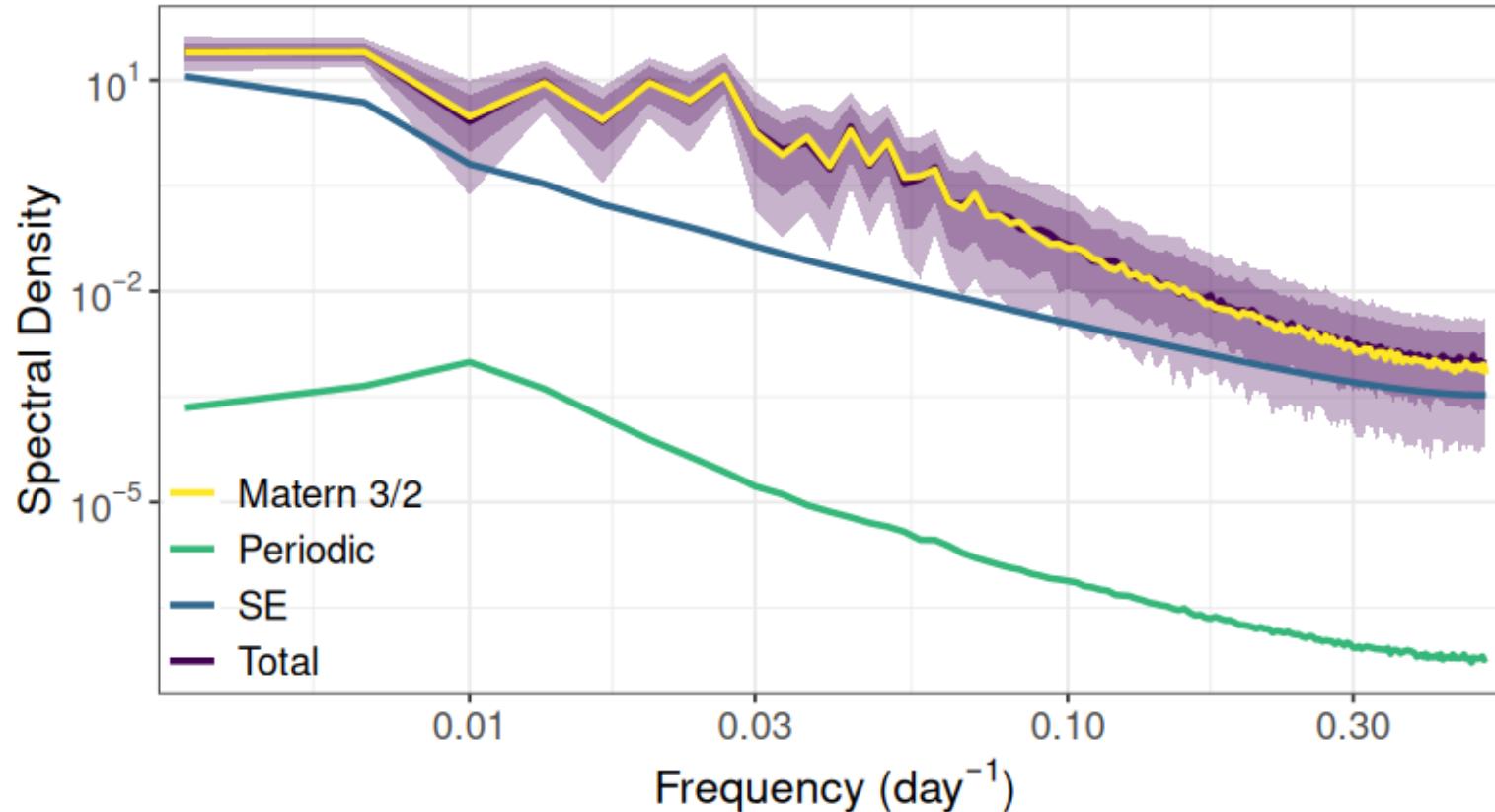
$$\ell_{\text{M32}} = 12.5$$

$$\sigma_{\text{P}} = 0.45$$

$$\ell_{\text{P}} = 37.6$$

$$T = 85.6$$

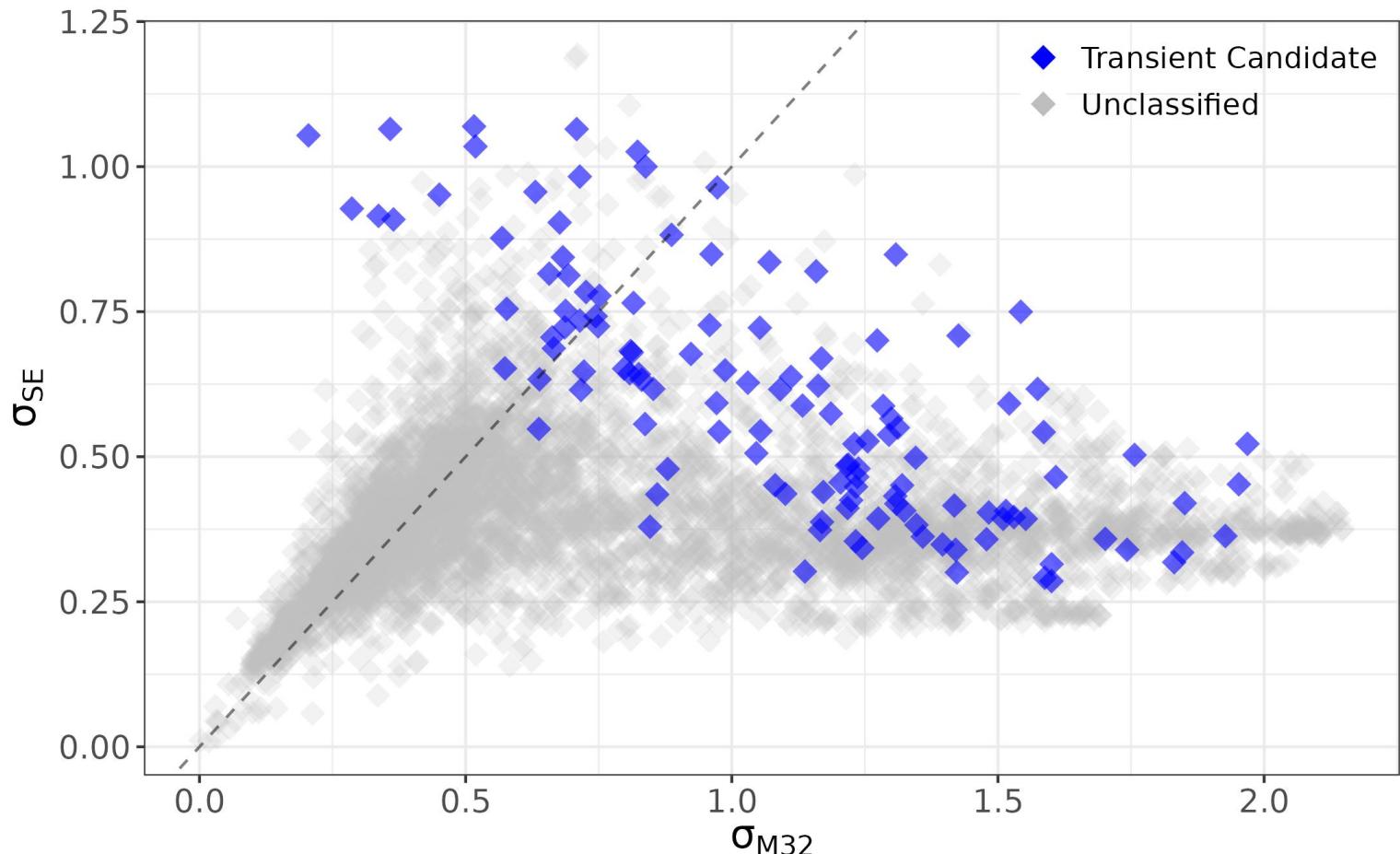
Power Spectral Density (PSD)



- Matern term dominates
- Very weak periodic term

Interpreting Amplitude as Transience

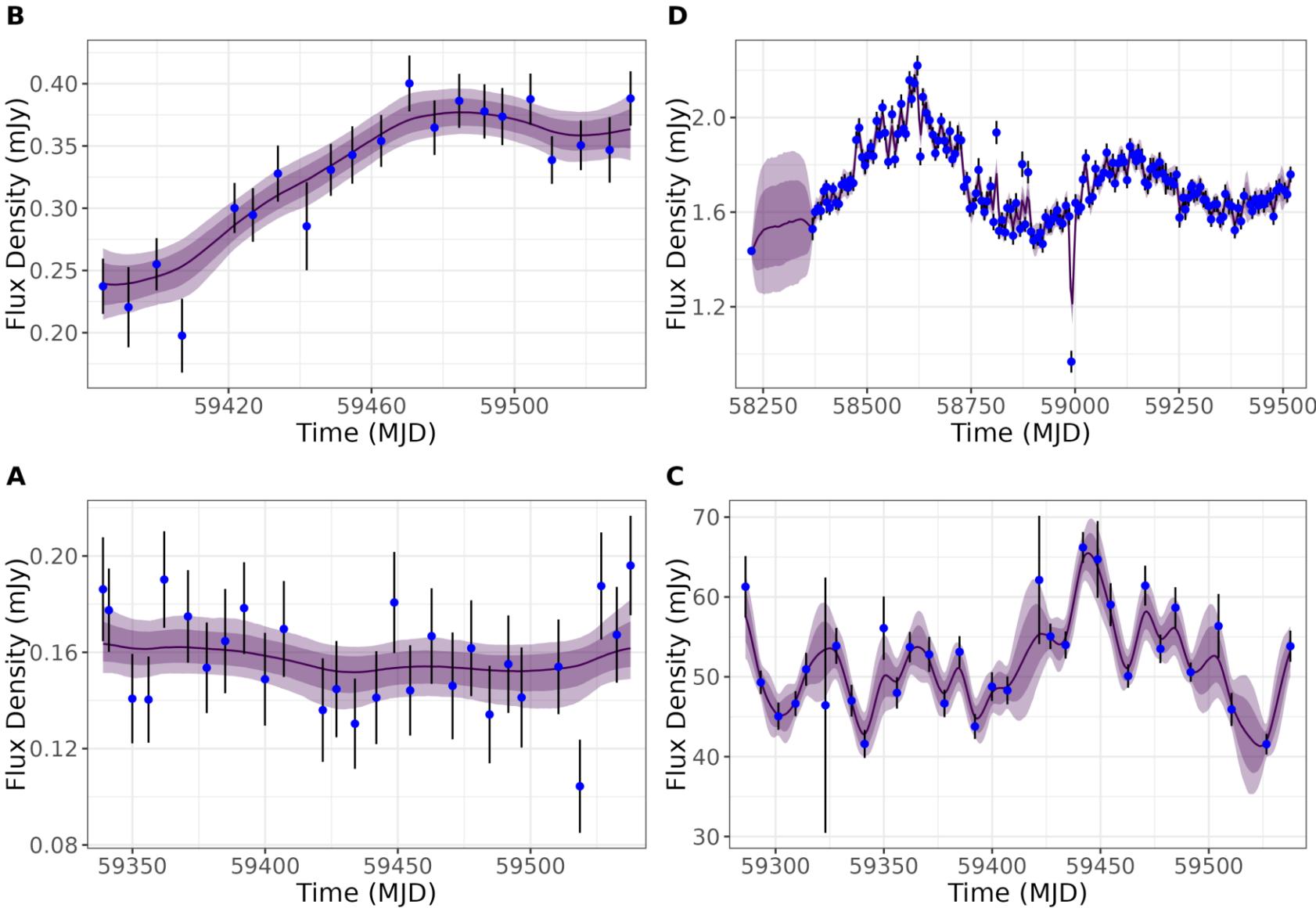
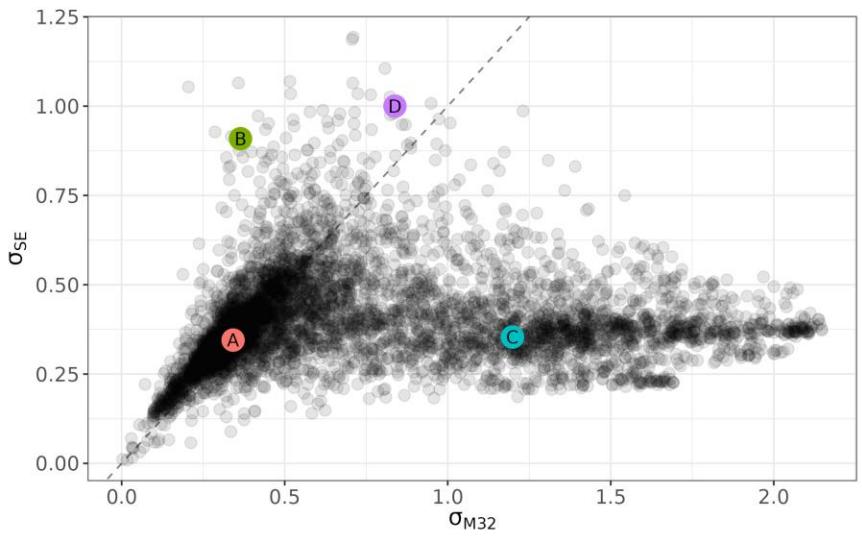
- Transience seems to manifest as large values in **amplitude**, σ .
- Previously identified transient candidates all seem to lie the upper right of this parameter space.



Data: Andersson et al. (2023)

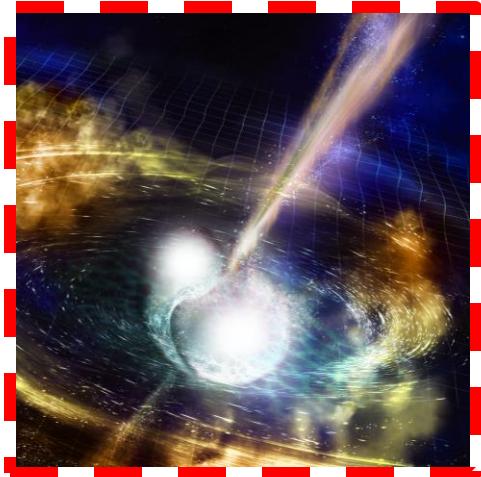
Figure: Fu et al. (in prep.)

Explore the hyperparameter space





Twinkle twinkle little star...



Exotic
phenomena



Large-scale
survey

Raw Data
Processing

Identify

Classify

10^3 to 10^6
light curves

Transient
candidates

Black holes,
supernova,
eclipsing
binary, GRB,
FRB, AGN,
etc, ...

... a Gaussian Process is what you are!



Tools

- Implemented in **Python¹** (v3.10) and **PyMC²** (v3.5.2)
 - Accessible to astronomers
 - Probabilistic programming framework
 - Well-maintained open-source software
- Repeated analyses in **R³** (v4.3.1) and **Stan⁴** (v2.34)
- Also considered: **celerite2⁵**, **george⁶**.

1. <https://www.python.org>

2. <https://www.pymc.io>

3. <https://cran.r-project.org/>

4. <https://mc-stan.org/>

5. <https://celerite2.readthedocs.io/en/latest/>

6. <https://george.readthedocs.io/en/latest/>